



# Global maps of Forel–Ule index, hue angle and Secchi disk depth derived from 21 years of monthly ESA Ocean Colour Climate Change Initiative data

Jaime Pitarch<sup>1,2</sup>, Marco Bellacicco<sup>3</sup>, Salvatore Marullo<sup>1,3</sup>, and Hendrik J. van der Woerd<sup>4</sup>

<sup>1</sup>Consiglio Nazionale delle Ricerche (CNR), Istituto di Scienze Marine (ISMAR),  
Via del Fosso del Cavaliere 100, 00133 Rome, Italy

<sup>2</sup>NIOZ Royal Netherlands Institute for Sea Research, Department of Coastal Systems, and Utrecht University,  
P.O. Box 59, 1790AB Den Burg, Texel, the Netherlands

<sup>3</sup>Italian National Agency for New Technologies, Energy and Sustainable Economic Development (ENEA),  
Via E. Fermi 45, 00044 Frascati, Italy

<sup>4</sup>Institute for Environmental Studies (IVM), Water & Climate Risk, Vrije Universiteit Amsterdam,  
De Boelelaan 1111, 1081HV Amsterdam, the Netherlands

**Correspondence:** Jaime Pitarch (jaime.pitarchportero@artov.ismar.cnr.it)

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**Abstract.** We document the development and public release of a new dataset (1997–2018), consisting of global maps of the Forel–Ule index, hue angle and Secchi disk depth. Source data come from the European Space Agency (ESA) Ocean Colour (OC) Climate Change Initiative (CCI), which is providing merged multi-sensor data from the mid-resolution sensors in operation at a specific time from 1997 to the present day. Multi-sensor satellite datasets are advantageous tools for ecological studies because they increase the probabilities of cloud-free data over a given region as data from multiple satellites whose overpass times differ by a few hours are combined. Moreover, data-merging from heritage and present satellites can expand the duration of the time series indefinitely, which allows the calculation of significant trends. Additionally, data are remapped consistently and analysis-ready for scientists. Also, the products described in this article have the exclusive advantage of being linkable to in situ historic observations and thus enabling the construction of very long time series. Monthly data are presented at a spatial resolution of  $\sim 4$  km at the Equator and are available at PANGAEA (<https://doi.org/10.1594/PANGAEA.904266>; Pitarch et al., 2019a). Two smaller and easier-to-handle test datasets have been produced from the former: a global dataset at  $1^\circ$  spatial resolution and another one for the North Atlantic at  $0.25^\circ$  resolution. The computer code for the generation of the Forel–Ule index, hue angle and Secchi disk depth from a given remote-sensing reflectance is also shared at <https://doi.org/10.5281/zenodo.4439646> (Pitarch et al., 2021) and can be easily set in loop mode for batch calculations.

## 1 Introduction

Single-sensor satellite datasets are not long enough to provide significant evidence of climatic trends that become manifest over fluctuations, so the multi-sensor approach has been fostered since the last decade (IOCCG, 2007). In Europe, the European Space Agency (ESA) Climate Change Initiative

(CCI) was created to address the United Nations Framework Convention on Climate Change requirements on a systematic monitoring of the global climate system. The Global Climate Observing System (GCOS) defined a number of essential climate variables (ECVs), which are physical, chemical or biological variables that critically contribute to the characterization of Earth's climate. Ocean Colour (OC) is one of the

ECVs, and the OC-CCI is devoted to producing a time series of consistent measurements encompassing water-leaving radiance in the visible domain, derived chlorophyll and inherent optical properties based on merged data from the MEdium Resolution Imaging Spectrometer (MERIS) (2002–2012), the Moderate Resolution Imaging Spectroradiometer (MODIS) (2002–present), the Sea-Viewing Wide Field-of-View Sensor (SeaWiFS) (1997–2010) and the Visible Infrared Imaging Radiometer Suite (VIIRS) (2012–present) sensors (Jackson et al., 2019).

While the chlorophyll concentration has historically been the most studied variable from ocean colour observations, more variables can be derived with confidence that can provide further insights, such as inherent optical properties or the diffuse attenuation coefficient ( $K_d(\lambda)$ ;  $\text{m}^{-1}$ ). Algorithm developers can also download the remote-sensing reflectance ( $R_{rs}(\lambda)$ ;  $\text{sr}^{-1}$ ) at <ftp://ftp.oceancolor.org/> (last access: 9 September 2019) and derive any other geophysical property themselves.

In the last decade, two historic optical variables have received renewed attention: the Secchi disk depth ( $z_{SD}$ ; m) and the Forel–Ule (FU; unitless) index. These two are the oldest oceanographic variables that are directly related to ocean colour, recorded for many decades before the advent of satellites (Boyce et al., 2012) and publicly available globally (NOAA, 2018). The Secchi disk is a white disk that is lowered from above the water surface and tracked visually until it goes out of sight. The depth when it ceases to be visible provides a reading of water transparency (Wernand, 2010). The  $z_{SD}$  is influenced by the underwater light attenuation, which in turn depends on the light absorption and scattering caused by the varying concentrations of dissolved and suspended substances in the water. The  $z_{SD}$  observation also depends on the disk design and environmental factors that need to be understood (Pitarch, 2020). The popularity and simplicity of this practice allowed its generalized use in oceanic surveys during the last century, although its use significantly declined in the last decades. Currently, the method is experiencing a revival due to recent efforts to derive this quantity from remote-sensing data (Lee et al., 2015).

The FU scale was developed as a set of standard colours that would allow visual water colour indexing (Wernand and Van Der Woerd, 2010). As with  $z_{SD}$ , the colour of natural waters is also influenced by the substances in the water, present in different concentrations, which constitutes the basis for ocean colour science. The popularity of the FU scale has had a similar history as the Secchi disk and has evolved from being widely adopted, then mostly abandoned and recently revived, aided by the link to remote-sensing data (Wernand et al., 2013).

The hue angle is a popular quantity in the field of image colour analysis. It is used as an intermediate quantity in the process of deriving FU from any light spectrum (Wernand et al., 2013), and it can be regarded as the continuous companion of the FU scale. Specific algorithms have been developed

for the spectral characteristics of various ocean colour sensors (Van Der Woerd and Wernand, 2015, 2018). In terms of in situ sampling, this quantity cannot be estimated visually, only radiometrically, although red–green–blue (RGB) images have proven good for this sake, thus enabling the use of digital cameras for citizen science (Ceccaroni et al., 2020).

Pitarch et al. (2019b) made an analysis of global seasonal variability in these parameters using the climatological ESA-OC-CCI data, including some cross-relationships and ranges of variability.

Despite the demonstrated usefulness of these optical parameters in marine physics and biogeochemistry, no public dataset of them exists thus far, so scientists need to calculate them from  $R_{rs}$  privately, which, other than creating redundant efforts, may be discouraging for non-experts. This gap is filled in the present article with the provision of the monthly series from 1997 to 2018 of Secchi disk depth, hue angle and Forel–Ule index at  $\sim 4$  km resolution.

## 2 Methods

### 2.1 Ingested data

The source product for all derived variables presented in this dataset is the merged multi-sensor OC-CCI v4.2  $R_{rs}$ , which is the primary quantity used for ocean colour studies, containing only spectral information related to the colour of the water while having all interfering factors ideally removed in the atmospheric correction and normalization calculations.  $R_{rs}$  is provided at a monthly frequency and projected on a rectangular grid of 2.5 arcmin for both latitude and longitude, which corresponds to about 4 km at the Equator, decreasing poleward. For more specific details, one can read the product user guide (Jackson et al., 2019). Numeric precision is single (32-bit floating-point values), except time, which is a 32-bit signed integer. Every downloaded data file contains a map of estimated bias in  $R_{rs}$ , which we add to generate the unbiased  $R_{rs}$  estimate. This bias compensation, which users need to perform themselves, has proven to enhance the quality of retrievals of  $R_{rs}$ -derived geophysical products like optical particle backscattering (Pitarch et al., 2020). Every generated file of Secchi disk depth, hue angle and Forel–Ule index keeps the same format as the corresponding source  $R_{rs}$  file.

### 2.2 The Forel–Ule index and hue angle algorithms

The hue angle ( $\alpha$ ;  $^\circ$ ) is a continuous numerical variable that expresses the colour of a light spectrum in a single number. In natural waters,  $\alpha$  ranges from about  $40^\circ$  in brown waters to about  $235^\circ$  in deep-blue oceanic waters. As the  $\alpha$  definition involves three spectral integrals across the continuous spectral range, some kind of adaptation needs to be done when applying it to satellite data that are provided only at a small number of wavelengths. Full technical details about  $\alpha$  de-

termination from multispectral  $R_{rs}$  are published (Van Der Woerd and Wernand, 2015, 2018; Pitarch et al., 2019b).

The FU scale was optically characterized as a canonical set of points  $(x, y)$  in the Commission Internationale de l'Éclairage (CIE) space (Novoa et al., 2013). Because  $\alpha$  is the polar angle of the  $(x, y)$  coordinates, the  $\alpha$  of all canonical FU points can be calculated. Thus, given an  $R_{rs}$  spectrum, its  $\alpha$  can be calculated, and subsequently, an FU index can be assigned to it by choosing the nearest FU index in  $\alpha$  units. In fact, FU can be seen as a discrete version of  $\alpha$  and can be used to cluster water masses in terms of their colour. However, it was found that the FU scale lacks resolution and dynamic range for the most oligotrophic oceanic waters (Pitarch et al., 2019b). Vast regions of the ocean remain saturated all year round at  $FU = 1$ . For this reason, a new lower end  $FU = 0$  was proposed in order to introduce more differentiation within the bluest waters. This dataset includes that addition.

### 2.3 The Secchi disk depth algorithm

We apply the state-of-the-art algorithm for the remote estimation of the Secchi disk depth ( $z_{SD}$ ) by Lee et al. (2015). According to the underlying theory,  $z_{SD}$  essentially depends on the spectral minimum of the diffuse attenuation coefficient ( $K_d$ ), which is analytically modelled as a function of the absorption ( $a$ ) and optical backscattering coefficient ( $b_b$ ) (Lee et al., 2013). The latter two are retrieved with the QAAv6 algorithm (Lee et al., 2002), including a correction for Raman scattering after Lee et al. (2013). It has been shown that such a step improves the match of the retrieved  $b_b$  against reference data (Pitarch et al., 2020). Solar zenith angle, which is also needed for  $K_d$ , is set equal to 0 to be consistent with the reflectance normalization in the OC-CCI dataset.

## 3 Product description

The processing is applied on a pixel basis from every source netCDF  $R_{rs}$  file. A sample code that applies the former algorithms to generate the hue angle, Forel–Ule index and Secchi disk depth from a given  $R_{rs}$  can be downloaded from <https://doi.org/10.5281/zenodo.4439646> and cited (Pitarch et al., 2021). This code can be easily set in loop mode and parallelized as well as linked to a netCDF file generator code, which is how the files of this dataset were produced.

The format of the product file is the same, and the spatial and temporal variables are directly copied from the source file. The hue angle, the Forel–Ule index and the Secchi disk depth are added. Content of each file is summarized in Table 1. Every netCDF file has the naming ESACCI-OC-L3S-MERGED-1M\_MONTHLY\_4km\_GEO\_PML\_Hue\_FU\_SD-yyyymm-fv4.0.nc, where yyyy is a four-character string for the year, and mm is a two-character string for the month.

From this product, two smaller, reduced-resolution products were generated:

1. a global dataset at  $1^\circ$  spatial resolution, more than enough to detect spatial features for global analyses;
2. a North Atlantic ( $0\text{--}60^\circ$  N,  $-80$  to  $0^\circ$  E) dataset at  $0.25^\circ$  resolution that can be used to study regional trends in this region at an increased resolution yet keeping the computational cost reasonably low.

## 4 Validation

In the remote-sensing community, validation is mostly known as a comparison to in situ reference data. Assuming errorless reference data, uncertainties are then due to the atmospheric correction and the applied water quality algorithms, though in situ data also contain uncertainties that are seldom considered.

Monthly OC-CCI products cannot be used for matchup to in situ data as matchups are intended for instantaneous data, with time coincidence between the satellite and the in situ readings recommended to be within a few hours. Understanding this requirement in a broad sense, daily satellite data, also provided within OC-CCI, may be acceptable. Therefore, uncertainties can be assessed using daily data, knowing that uncertainties present in the monthly aggregates will be smaller as the temporal averaging will dampen random noise by a factor equal to the square root of the number of observations at every pixel for every month, and, in fact, reduced noise can be seen in the produced monthly images, where speckle noise looks rather low. Systematic biases will remain after averaging, though.

Ideally, uncertainties would be provided at pixel level for every product map, but such information could only be given after an uncertainty differentiation across the product dynamic range, using a large and high-quality in situ dataset that spans the broadest dynamic range, followed by a clustering into spectral classes (Jackson et al., 2019). In fact, such a stringent requirement means that, for some products like backscattering, uncertainties cannot be provided as maps. As the satellite variables derived in this article are also affected by a lack of open and quality-controlled in situ data, uncertainties are provided as bulk estimates.

### 4.1 Remote-sensing reflectance

The OC-CCI's product user guide (PUG) (Jackson et al., 2019) documents uncertainty assessment of satellite OC-CCI  $R_{rs}$  by comparing it to a large in situ  $R_{rs}$  database that merges in situ  $R_{rs}$  from many different sources with very disparate instrumentation and quality standards (Valente et al., 2019) without an uncertainty budget assessment. Usage of this entire database can artificially increase the error statistics when compared to satellite data. For this reason, in a previous publication (Table 3 in Pitarch et al., 2020) we compared

**Table 1.** File description.

Parameter	Description	Dimensions	Size (pixels)
lat	Geographic latitude, positive northward	Degree	4320 × 1
long	Geographic longitude, positive eastward	Degree	8640 × 1
time	Time since 1 January 1970 00:00:00 UTC	Day	1 × 1
alfa	Hue angle according to Pitarch et al. (2019b)	Degree	8640 × 4320 × 1
FU	Extended Forel–Ule index of the remote-sensing reflectance (0 to 21) according to Pitarch et al. (2019b)	–	8640 × 4320 × 1
zSD	Secchi disk depth, according to Lee et al. (2015)	m	8640 × 4320 × 1

daily OC-CCI reflectances to Valente’s database as well as to two other internal databases. Uncertainties were the lowest when OC-CCI daily products were compared to an internal database “CNR”, collected using strict NASA protocols and uncertainties traceable to National Institute of Standards and Technology (NIST) standards. Relative RMSEs were within 20 %, with biases not significantly different from 0, except for the band at 412 nm. At 670 nm, errors were much higher in relative terms due to the low value of  $R_{rs}(670)$ , but they remained low in absolute units ( $sr^{-1}$ ). These RMSEs include uncertainties present in the in situ  $R_{rs}$  as well, which may range between 5 % and 10 % (Zibordi et al., 2019), so an educated guess of a spectral average of 7.5 % was subtracted in quadrature from them to isolate the errors in the satellite data.

These remaining uncertainty estimates need to be propagated by the Secchi disk depth, hue angle and Forel–Ule index algorithms. To assess this matter, three example  $R_{rs}$  spectra were taken, representing ultra-oligotrophic, oligotrophic and eutrophic waters. Their respective Secchi disk depth, hue angle and Forel–Ule index were calculated. Posteriorly, a Monte Carlo analysis was made by adding the  $R_{rs}$  uncertainties as Gaussian-independent noise. This was repeated 10 000 times, and the histograms (Fig. 1) and related statistics (Table 2) of the resulting Secchi disk depth, hue angle and Forel–Ule index were retrieved. It is shown that, for the Secchi disk depth, biases have a spread of about  $\sigma = 22\%$  around the errorless value, with a negligible mean bias, increasing towards more turbid waters. The hue angle experiences a spread of about  $\sigma = 1\%$ , also with a negligible bias, though this result has to be interpreted keeping in mind the high average value and the low dynamic range of the hue angle. Statistics of the FU are much harder to interpret due to the strongly discrete nature of this variable, but essentially FU oscillates between the central value and the adjacent ones, in proportions that depend on the distances to these in hue angle terms.

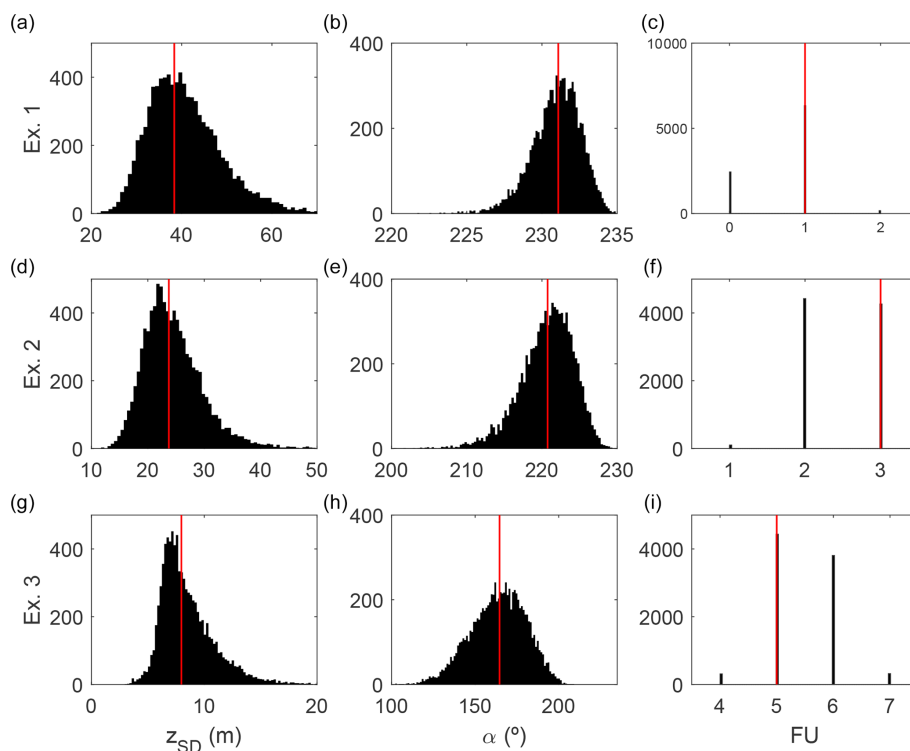
A third uncertainty source is the remote algorithm itself with respect to the actual in situ value. Since the algorithms were published elsewhere, this article does not validate them further. Here, we simply summarize statistical findings on their accuracy in each of the next three sub-sections.

## 4.2 Secchi disk depth

The Secchi disk depth algorithm was validated by Lee et al. (2015). The  $z_{sd}$  derived from in situ  $R_{rs}$  compared well to actual measurements with a Secchi disk, with a mean absolute percent difference (MAPD) of  $\sim 19\%$ . The linear regression slope between in situ and remote estimation deviated marginally from the 1 : 1 line. Unfortunately, most of their data were not open to external users, and the MAPD is a statistical parameter different from rms or standard deviation. However, a quick numerical verification with Gaussian noise (not shown) reveals that the ratio between the relative RMSE and MAPD for unbiased estimates is very stable at about  $\sim 1.25$ . Therefore, Lee’s MAPD = 18.2 % equals RMSE = 22.8 %. From this uncertainty, one must subtract the part due to the in situ radiometry. Unfortunately, in Lee’s data, radiometric uncertainties are not traced, so an educated guess of an average 10 % RMSE is made here. This guess is subtracted in quadrature from the above 22.8 %, which results in an uncertainty of  $\sim 20\%$  for Secchi disk depth model estimations with respect to in situ data. Overall, adding in quadrature the uncertainties due to the satellite radiometry, we may conclude that Secchi disk depth derived from daily OC-CCI data has an uncertainty of  $\sim 32\%$  when compared to in situ data.

## 4.3 Hue angle

Uncertainties in the  $\alpha$  product are derived from the approximation of the spectral integrals involved in the algorithm from a limited set (six here) of satellite bands (Van Der Werd and Wernand, 2015, 2018; Pitarch et al., 2019b). These uncertainties vary with the water type because the few available spectral bands may be enough to capture all the optical information in simple case 1 waters but not in optically complex waters. In fact, for case 1 waters, in relative terms, this amounts to less than 0.1 %, while in optically complex waters, it can amount to up to 3 %. We can therefore conclude that, for satellite  $\alpha$  estimations, the uncertainty is almost totally caused by the satellite radiometry, which ranges from less than 1 % to about 11 % as water shifts from very clear to turbid. This may also be an upper bound because we have assumed spectrally independent Gaussian noise at every satel-



**Figure 1.** Monte Carlo simulation of error propagation from OC-CCI  $R_{rs}$  to Secchi disk depth, hue angle and Forel–Ule index. Panels (a)–(c) show the results of 10 000 simulations (black bars) for ultra-oligotrophic waters compared to the reference spectrum results (red line). Panels (d)–(f) show the results for an oligotrophic water spectrum, and (g–i) depict the results for a eutrophic water spectrum.

**Table 2.** Biases and standard deviation ( $\sigma$ ) of the errors caused by random additive Gaussian perturbations to three given remote-sensing reflectances from which the Secchi disk depth, the hue angle and the Forel–Ule index are derived. The “exact values” correspond to the red vertical lines in Fig. 1.

$z_{SD}$ (m)			$\alpha$ ( $^{\circ}$ )			FU		
Exact value	Bias (%)	$\sigma$ (%)	Exact value	Bias (%)	$\sigma$ (%)	Exact value	Bias (%)	$\sigma$ (%)
38.4	6.9	22.5	231	−0.06	0.7	1	−25.4 %	47.7 %
28.5	3.3	22.7	226	−0.05	1.6	3	−17.1 %	18.0 %
8.0	6.6	30.3	165	−0.5	10.0	5	9.3 %	12.6 %

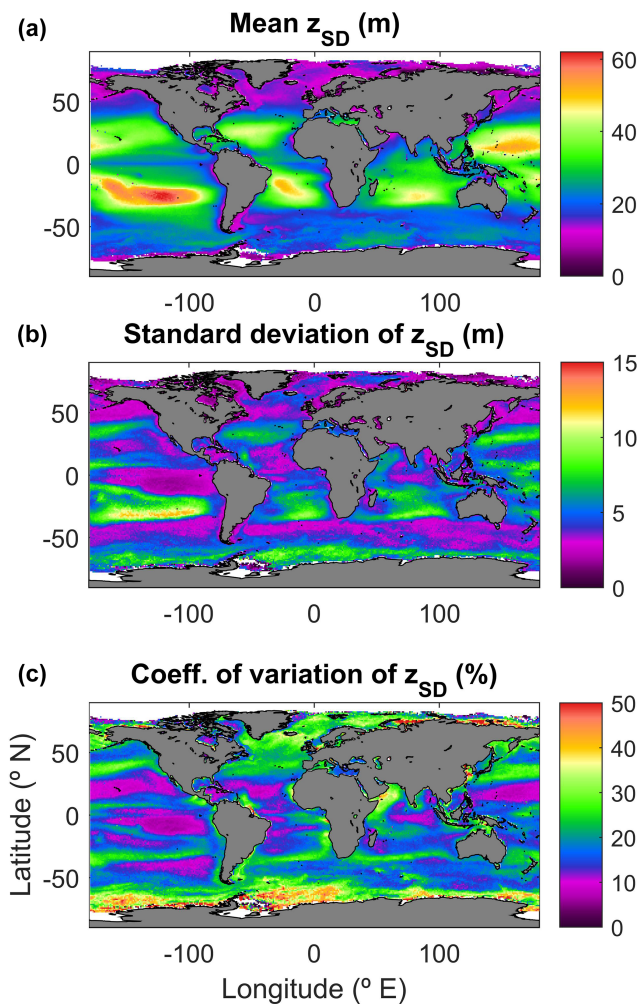
lite band, but errors in an atmospheric correction may have some spectral correlation. As a consequence of this, satellite-derived hue angle (and FU) seem to be particularly robust to uncertainties in satellite  $R_{rs}$  after atmospheric correction (Wang et al., 2018).

#### 4.4 Forel–Ule index

In situ FU readings are different depending on whether a submerged Secchi disk is used to obtain a visual estimate of the colour over its surface or whether the estimate is obtained by looking at the water only. Essentially, an underwater white object reflects light that, after propagating back across the water column, is observed from above with a hue angle that is lower than that of the surrounding water. These

differences are significant, and the amount of this bias depends on the colour of the water effect itself, which was previously shown after radiative-transfer simulations (Pitarch, 2017) and recently shown experimentally (Nie et al., 2020). In addition, in situ FU can either be estimated visually or derived mathematically from an  $R_{rs}$  that has been measured in situ. The equivalence of both must not be taken for granted, and their relation has not been studied yet. In summary, there are several methodological concerns that affect the definition of reference FU measurements for comparison to satellite estimations.

In the interim, we choose to define reference FU measurements as those originating from in situ hyperspectral  $R_{rs}$ , of which the hue angle has been calculated by integration and FU by discretization. In such a case, the uncertainties

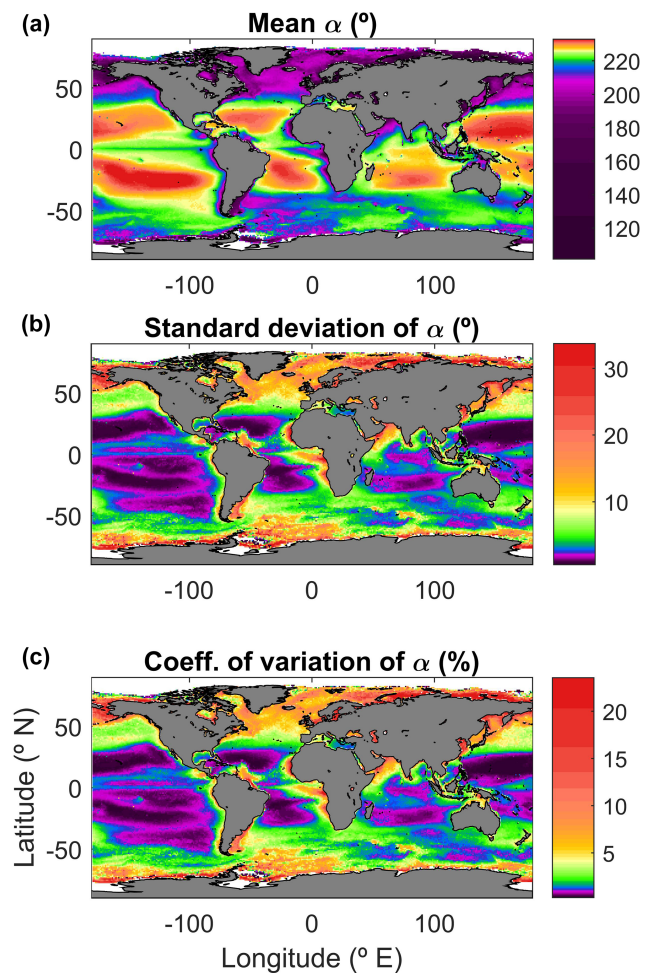


**Figure 2.** Global maps of the Secchi disk depth ( $z_{SD}$ ) in the period 1997–2018. (a) Series mean. (b) Standard deviation. (c) Coefficient of variation.

are only those related to the satellite radiometry. The error is bounded within  $\pm 1$  in almost all cases.

## 5 Global-average maps

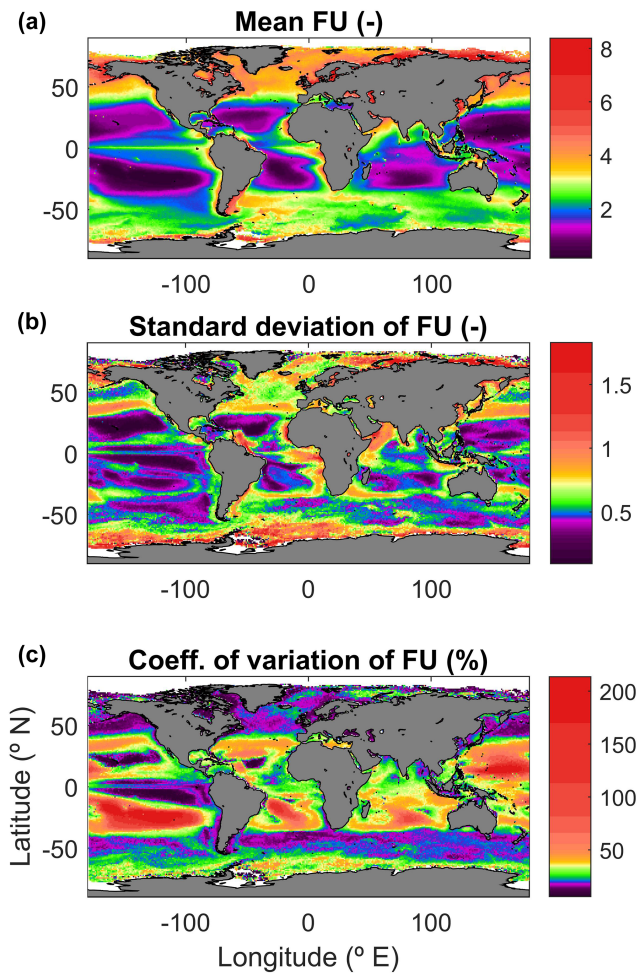
In order to provide the reader with a feeling of the global variability covered by these parameters, some basic statistics are demonstrated in this section. These results were produced using the reduced-resolution global product at  $1^\circ$ . At every pixel, the 256 values (21 years and 4 months of data) were combined, and the arithmetic mean, standard deviation and coefficient of variation were calculated. Figure 2 shows the results for the  $z_{SD}$  time series. The mean map shows a global pattern that is similar when deriving other parameters like the surface chlorophyll-*a* concentration. In the oligotrophic gyres, there is high transparency induced by low concentration of phytoplankton cells, as also highlighted by the  $z_{SD}$ . For example, the South Pacific subtropical gyre reaches a



**Figure 3.** Global maps of the hue angle ( $\alpha$ ) in the period 1997–2018. (a) Series mean. (b) Standard deviation. (c) Coefficient of variation.

mean of  $\sim 60$  m in its most transparent core. The standard deviation of  $z_{SD}$  shows interesting patterns. It is mostly driven by the seasonal variability, and it is highest in the gyre boundaries, showing the seasonal expanding and shrinking of the subtropical gyres, though they might also include the long-term effect of gyre expansion (Polovina et al., 2008). Other peaks of variability are zones affected by seasonal blooms like off Iceland or the Antarctica coast. This latter feature is most accentuated in the coefficient of variation.

The hue angle ( $\alpha$ ) shows a pattern similar to  $z_{SD}$  (Fig. 3), though with a marked non-linearity. Global values cover the range  $\sim 50$ – $235^\circ$ , though 85 % of the world's oceans are blue with  $\alpha > 200^\circ$ . The standard deviation is very low in the oligotrophic areas, reflecting the fact that the human eye has little sensitivity to variations in colour when these happen at the lower spectral end of the visible range. On the other hand, variability is mostly highlighted in zones of green waters like mesotrophic and eutrophic areas as well as turbid shelf seas and coastal zones.

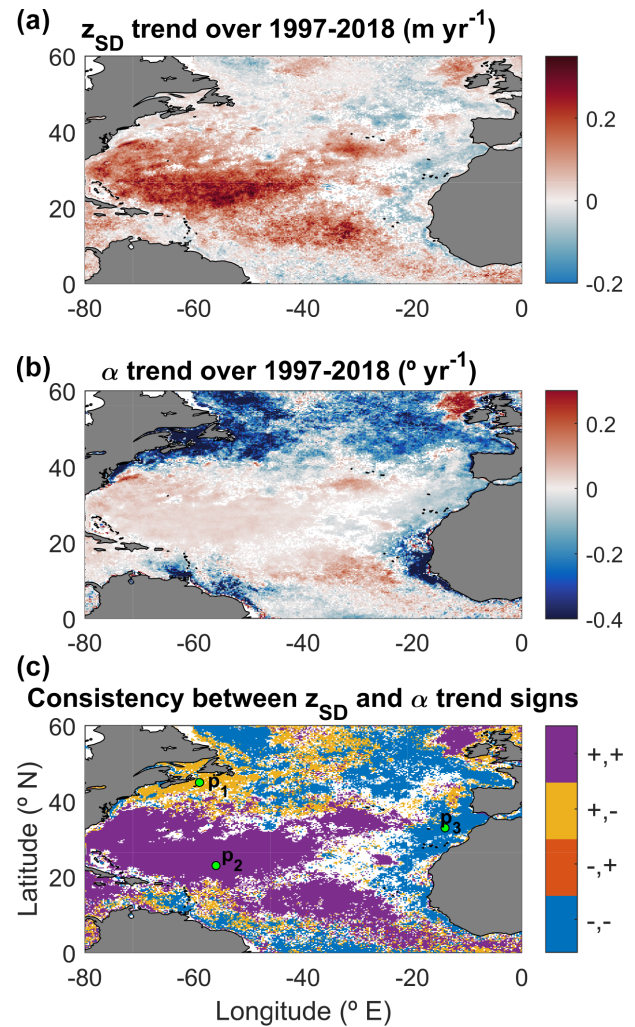


**Figure 4.** Global maps of the Forel–Ule (FU) index in the period 1997–2018. (a) Series mean. (b) Standard deviation. (c) Coefficient of variation.

The Forel–Ule index mean value (Fig. 4) reflects the known global patterns. Despite ranging the scale from 0 to 21, 99 % of the data are within  $FU \leq 8$ , keeping the upper rest of the FU scale for very turbid coastal areas or highly humic waters. Note that, despite the fractional results due to statistics, FU are integers. As shown for the hue angle, FU shows very little variation for the oligotrophic and low-latitude areas, reaching the widest dynamic range in coastal and high-latitude blooming seas.

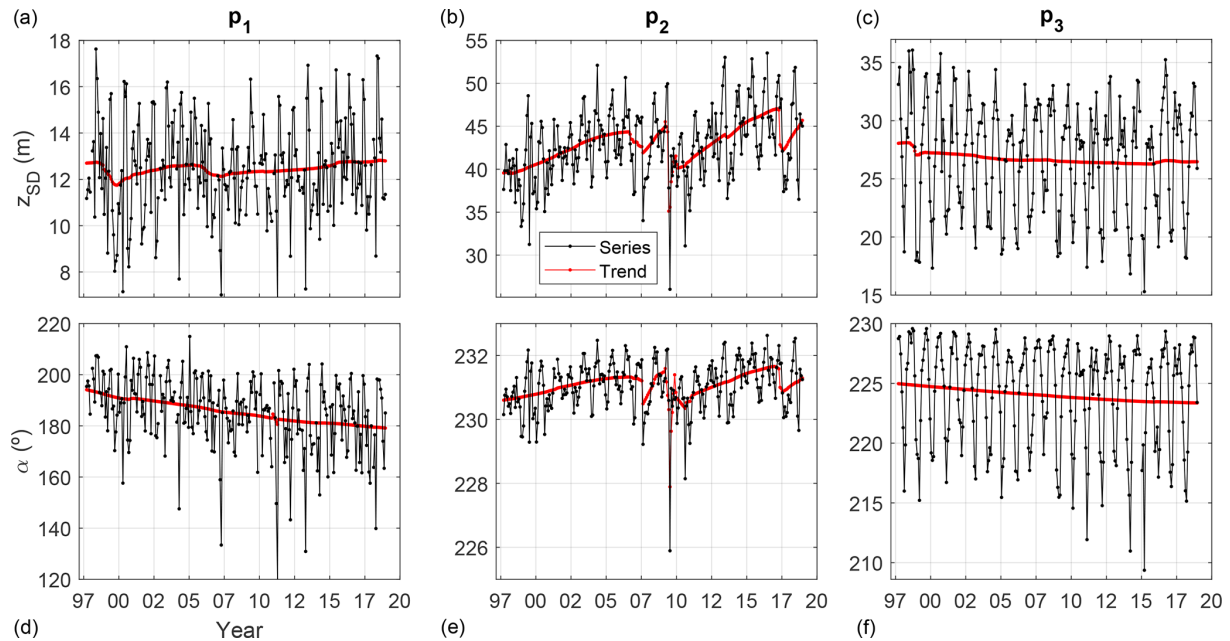
### 6 Example of application: trend analysis for the North Atlantic Ocean

This section provides an example of application that shows the unique possibilities of long-term satellite data: the calculation of pixel-wise oceanic trends. Here, we used the lower-spatial-resolution product at  $0.25^\circ$ , covering the Atlantic Ocean north of the Equator. This section reports a trend analysis of the area for the Secchi disk depth and the hue



**Figure 5.** (a) Linear slope of Secchi disk depth ( $z_{SD}$ ) trend at pixel scale for the North Atlantic. (b) Same as before but for the hue angle ( $\alpha$ ). (c) Signs of both linear slopes.

angle. Seasonal decomposition was made pixel-wise with the publicly available BEAST code (Zhao et al., 2019). The Sen’s slope (Sen, 1968) of the de-seasonalized signal was calculated, and the result was tested for significance at 95 % (Mann, 1945; Kendall, 1975). In Fig. 5, the upper two panels only show the statistically significant trend signals for  $z_{SD}$  and  $\alpha$ , calculated for the period 1997–2018. A positive (negative) trend in  $z_{SD}$  means that a Secchi disk will disappear from sight at a larger (shallower) depth, corresponding to a transparency increase (decrease). A positive (negative) trend in  $\alpha$  indicates a blue (red) shift, normally corresponding to a reduction (increase) in the surface chlorophyll-*a* concentration (Pitarch et al., 2019b). In the lowest panel of Fig. 5, the signs (positive or negative) of the trends are combined to yield four classes; a combination of positive  $z_{SD}$  sign and negative hue angle sign is given by the code (+−). Again, no colour indicates no significant trend in both indexes.



**Figure 6.** Time series of Secchi disk depth (a–c) and hue angle (d–f) at three points in the North Atlantic Ocean. Red lines show the de-seasonalized signal.

The North Atlantic trend maps (Fig. 5) show that the subtropical gyre is gaining transparency and getting bluer and bluer. On the other hand, mesotrophic, more northern waters outside the gyre as well as coastal waters are experiencing totally different dynamics. In general, there is a loss of transparency and a red shift. Some areas where trends are very small have apparent opposite trends in  $z_{SD}$  and  $\alpha$ , which highlights the fact that these two quantities do not completely contain the same information, and it will deserve some future research. In Fig. 6, the complexity in the spatial and temporal behaviour is detailed for three points, their position indicated on the lower map of Fig. 5. We anticipate that the study of these variations at various spatial scales and their relation to potential physical and environmental drivers like sea surface temperature, ocean heat content, nutrients, mixed-layer depth, coastal currents and river outflow will be explored.

## 7 Code and data availability

The 256 full-resolution netCDF files constituting the global time series at monthly frequency and 4 km resolution are publicly available through the following link: <https://doi.org/10.1594/PANGAEA.904266> (Pitarch et al., 2019a). The repository also contains two sub-products of much smaller size that can be quickly downloaded and easily handled locally: (1) a downgraded global dataset at  $1^\circ$  for latitude and longitude and (2) a downgraded dataset at  $0.25^\circ$  for latitude and longitude, for the North Atlantic Ocean.

The computer code for the generation of the Forel–Ule index, hue angle and Secchi disk depth from

a given remote-sensing reflectance is available at <https://doi.org/10.5281/zenodo.4439646> (Pitarch et al., 2021).

## 8 Conclusions

We have derived a consistently calibrated 21-year-long dataset of monthly Secchi disk depth, hue angle and Forel–Ule index over the global oceans at  $\sim 4$  km spatial resolution at the Equator. The algorithms employed for this release have been recently published and are at the state of the art, and their application is not restricted to any specific water type. Best-effort validation statistics have been provided based on current evidence. Global-average maps have shown meaningful patterns that are consistent with the general understanding of the ocean dynamics. A small example of trend estimation has also been provided to illustrate one of the main areas of research that can be addressed with this dataset.

This dataset has been directly derived from the OC-CCI dataset and, as such, shares its advantages and limitations. Main advantages are the careful sensor inter-calibration to avoid spurious trends, the constant revision of the methodology in search of improved quality and the warranted continuation of the time series for the next years. Limitations are the exclusion of high-latitude areas due to high solar zenith angle, which causes problems in the OC-CCI processing, and the uneven monthly binning depending on cloud cover. These issues could be solved in a future release, which could use daily source data with an application of a gap-filling procedure.

With this release, we hope to eliminate redundant efforts within the community to estimate water transparency and colour from satellite measurements. Depending on the reception and user needs, future releases using other temporal and spatial resolutions could be made.

The Secchi disk depth and the Forel–Ule index time series could be further enlarged backwards by linking them to the archived *in situ* data that go back to the year 1890 (Boyce et al., 2012). However, even in the absence of historic data, 21 years of satellite data are already a valuable dataset that may provide new insights into the ocean’s variability in relation to climate change and constitute valuable information to train and validate mechanistic oceanic models.

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