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Optimization of the LIBS Technique in Air, He, and Ar at Atmospheric Pressure for Hydrogen Isotope Detection on Tungsten Coatings

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Abstract

In current and future fusion devices, detecting hydrogen isotopes, particularly tritium and deuterium, implanted or redeposited on the surface of Plasma-Facing Components (PFCs) will be increasingly important to ensure safe machine operations. The Laser-Induced Breakdown Spectroscopy (LIBS) technique has proven capable of performing this task directly in situ, without handling or removing PFCs, thus limiting analysis times and increasing the machine's duty cycle. To increase sensitivity and the ability to discriminate between isotopes, LIBS analysis can be performed under different background gases at atmospheric pressure, such as air, He, and Ar. In this work, we present the results obtained on tungsten coatings enriched with deuterium and/or hydrogen as a deuterium–tritium nuclear fuel simulant, measured with the LIBS technique in air, He, and Ar at atmospheric pressure, and discuss the pros and cons of their use. The results obtained demonstrate that both He and Ar can improve the LIBS signal resolution of the hydrogen isotopes compared to air. However, using Ar has the additional advantage that the same procedure can also be used to detect He implanted in PFCs as a product of fusion reactions without any interference. Finally, the LIBS signal in an Ar atmosphere increases in terms of the signal-to-noise ratio (SNR), enabling the use of less energetic laser pulses to improve performance in depth profiling analyses.

Keywords: LIBS; hydrogen isotope detection; tungsten PFCs



Academic Editors: Jan Willem Coenen and Thomas P. Davis

Received: 6 May 2025

Revised: 10 June 2025

Accepted: 23 June 2025

Published: 1 July 2025

Citation: Almaviva, S.; Baiamonte, L.; Pistilli, M. Optimization of the LIBS Technique in Air, He, and Ar at Atmospheric Pressure for Hydrogen Isotope Detection on Tungsten Coatings. *J. Nucl. Eng.* **2025**, *6*, 22. <https://doi.org/10.3390/jne6030022>

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1. Introduction

Monitoring the accumulation of tritium (T) inside a vacuum vessel (VV) is of paramount importance in future fusion devices and crucial for ensuring the safe operation of future tokamaks. Once tritium retention exceeds a critical threshold (e.g., 700 g in ITER) [1], it may be necessary to interrupt plasma operations until the stored tritium can be removed. Tritium retention in the VV occurs through direct implantation into the surfaces of Plasma-Facing Components (PFCs) and the co-deposition of materials that are eroded from the surfaces of PFCs in some areas of the VV and transported to other areas [1–3]. Currently, tungsten (W) is a candidate material for ITER PFCs due to its properties, such as a high melting point, high thermal conductivity, low sputtering yield, and low retention rate for hydrogen isotopes, and it has therefore been proposed both as a surface coating for structural materials and as a bulk material for PFCs.

To date, research aimed at quantifying the T content in PFCs has been mainly supported by post mortem analyses [4], once removed from the VV during machine maintenance periods. In the future, however, to reduce these periods and increase the machine duty cycle, access to PFCs should be more limited. This need supports the development of complementary, in situ measurement methods that work without removing PFCs from the VV. Laser-Induced Breakdown Spectroscopy (LIBS) [5] is a candidate technique to solve these issues, attracting increasing interest for its features. LIBS can indeed perform in situ qualitative and quantitative analytical chemistry studies on PFCs [6–8], simultaneously revealing all the chemical elements and the possible presence of hydrogen isotopes without any preliminary sample preparation. LIBS typically requires only a few μg of sample for analysis, leaving any PFC substantially intact for any subsequent analysis, and can be applied both in vacuum [9,10] and at atmospheric pressure, both in contact and at several meters from the sample [9]. Finally, LIBS is sensitive to materials with a structure composed of multiple thin overlapping layers, with thicknesses of up to a few hundred nm [9,11], which makes it suitable for performing depth profiling analyses and investigating the fuel retention, erosion, or re-deposition of materials in PFCs.

In this work, we report on the LIBS analysis of fusion-relevant metallic coatings composed of W enriched with D, H, or both to simulate unburned nuclear fuel deposited or implanted on the surfaces of PFCs. The measurements were carried out under a He or Ar flux of 2 L/min to improve the spectral resolution of the Balmer alpha lines (D_α and H_α at 656.1 nm and 656.28 nm, respectively) of the two isotopes. The measurements were compared to similar standard measurements carried out in air. A high depth resolution, down to 500 nm per single laser shot, was obtained, while still maintaining a good SNR for spectral features ascribable to the emitting species. The application of the two background gases, He and Ar, led to the conclusion that He could be a good choice when the discrimination and spectral resolution of the D_α and H_α peaks are a priority in the investigation, whereas Ar is better for increasing the SNR of the whole signal and applying a reduced pulse energy to increase the depth resolution and detect other elements. A semi-quantitative estimation of the D-H contained in the samples was performed by applying the calibration-free (CF) method [12], which does not require any reference or calibration sample.

2. Materials and Methods

A sample simulating co-deposits of W with D and H was produced at the National Institute for Laser, Plasma and Radiation Physics in Romania (NILPRP) by High-Power Impulse Magnetron Sputtering (HiPIMS) [13]. It consists of a molybdenum (Mo) sample with an area of $11 \times 15 \text{ mm}^2$ and a 1 mm thick metal plate. The sample (W-D sample hereafter) is composed of a 5 μm thick W surface layer, co-deposited with D on a Mo substrate.

The LIBS system is composed of a Montfort M-NANO PIV Nd:YAG double-pulse laser (wavelength: 1064 nm; pulse duration: 8 ns; max repetition rate: 20 Hz; output spot diameter: about 3 mm) capable of delivering two laser pulses, with a maximum energy $\sim 65 \text{ mJ}$ each, temporally delayed up to 300 μs [7]. The two pulses were set to overlap (interpulse delay = 0 s) to act as a single pulse and were focused by a plano-convex lens on the target with a 25 mm diameter and $f = 100 \text{ mm}$. From the observation of the LIBS craters through an optical microscope, the diameter of the laser spots on the target's surface was found to be $\sim 350 \mu\text{m}$. The LIBS plasma was collected by a 50 mm diameter UVFS plano-convex lens, with $f = 150 \text{ mm}$, on an optical path of collection almost collinear to the laser beam, tilted by 10 degrees, to collect light mainly from the central region of the plasma and reduce the influence of the peripheral zones, which are known to have a lower temperature and electron density than the central part [14,15]. The lens directed the LIBS

signal towards a 2 mm diameter circular fiber bundle, composed of 17 fibers, N.A. 0.22 with a 245 μm diameter, plus a single fiber, N.A. 0.22 with a 50 μm diameter. The 6 m long bundle was bifurcated at 4 m into two branches: the first contained the 17 fibers embedded in an S.S. ferrule and aligned to create a 4.16 mm long linear bundle, while the second contained the single fiber in an SMA connector.

The two branches were connected to two different spectroscopic systems: the first branch to an ISA Jobin Yvon "TRIA550" Czerny–Turner monochromator, with a 2400 g/mm grating ($\lambda/\Delta\lambda = 50,000$) and a resolution of $\Delta\lambda = 656/50,000 \sim 0.013$ nm at 656 nm, sufficient to detect the nearby D_α and H_α lines. An Andor "ISTAR" ICCD camera DH534-18 F (1024 \times 256 pixels) recorded the LIBS signal, covering a spectral window of up to 12 nm. The second branch was connected to an Echelle "Aryelle 100" spectrometer by LTB, capable of recording the LIBS signal over a broad spectral range (225–760 nm) with a resolution of up to 9000. A second Andor "ISTAR" camera completed the setup. A schematic diagram and picture of the entire system are shown in Figure 1:

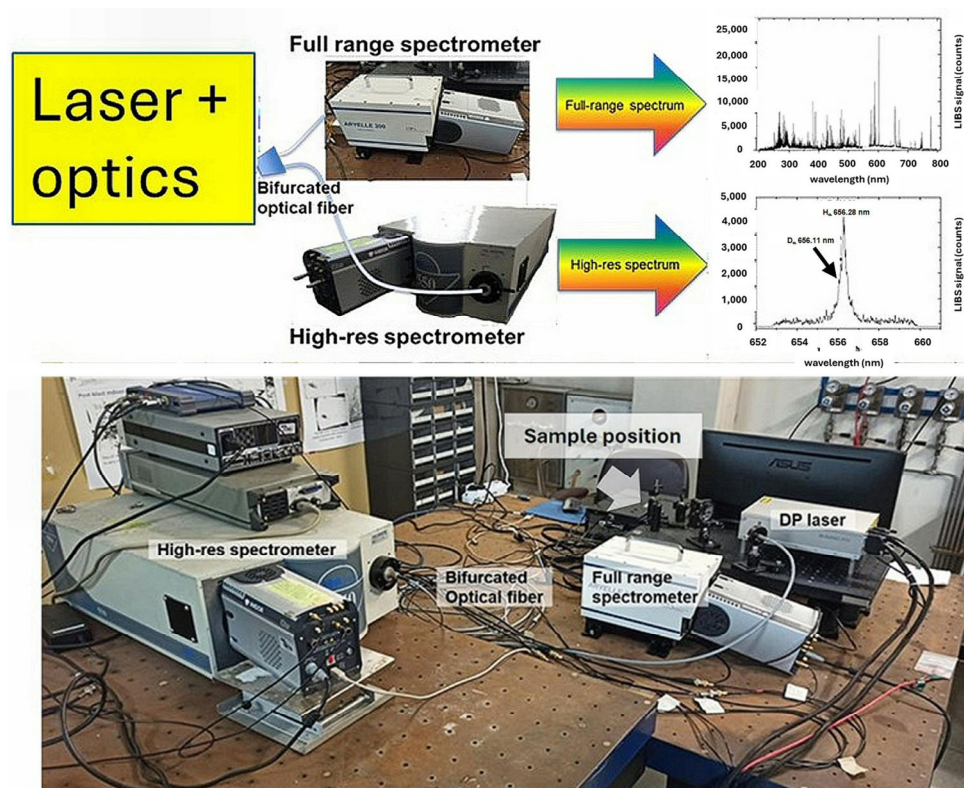


Figure 1. (top) Schematic diagram of the LIBS apparatus. (bottom) A picture of the system with the key elements.

3. Results and Discussion

3.1. Detection of D on the W-D Sample

The W-D sample was used to optimize the detection of D in the presence of the interfering H emission from the environment, using two background gases (He and Ar) and comparing the signal with respect to LIBS in air. For each gas, the pulse energy was set to 50 mJ, with a repetition rate of 2 Hz, while the gate width of the ICCD was fixed to 3 μs . The gate delay was optimized in the three cases to obtain the maximum resolution of the two D_α and H_α peaks while still maintaining a good SNR.

Figure 2 shows an example of the LIBS signal obtained by the Czerny–Turner spectrograph in the case of He (top), air (center), and Ar (bottom) as background gases for three different gate delays.

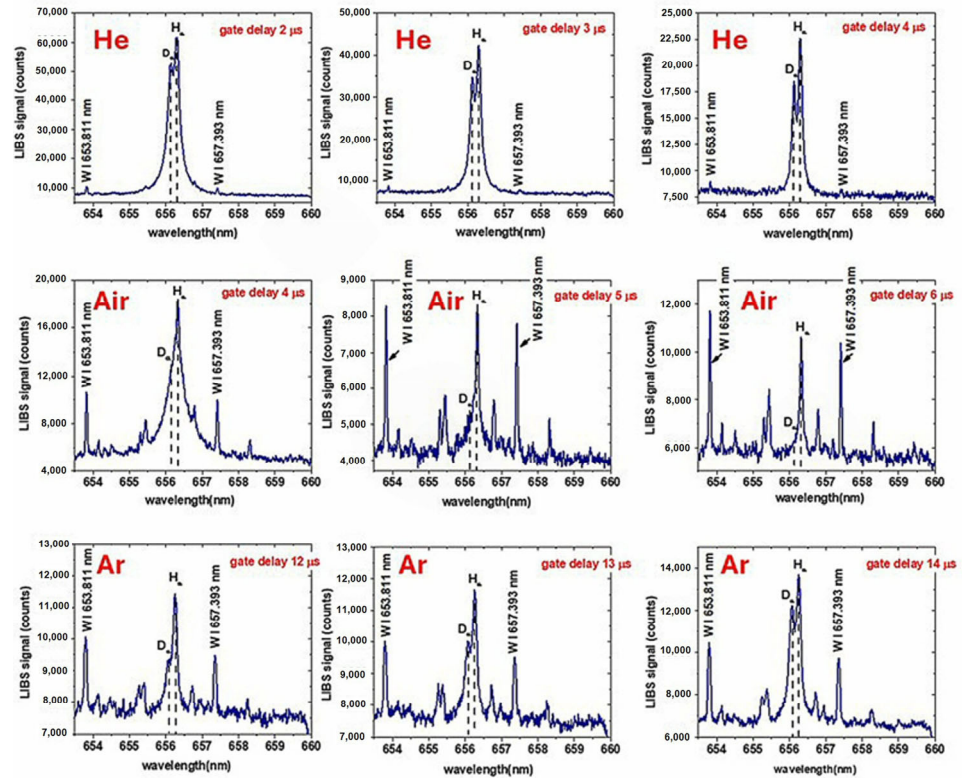


Figure 2. (Top) The LIBS signal in the Balmer alpha region of the spectrum of the W-D sample under He flux at delays of 2, 3, and 4 μs from the laser pulse; (Center) in air at delays of 4, 5, and 6 μs from the laser pulse; and (Bottom) under Ar flux at delays of 12, 13, and 14 μs from the laser pulse. For all the measurements, the laser pulse energy was set to 50 mJ and the ICCD gate width was set to 3 μs .

From Figure 2, it can be observed that using He or Ar as the background gas significantly enhances the spectral resolution of the two D_{α} and H_{α} peaks compared to the standard LIBS in air when the delay of the ICCD is appropriately set. In particular, the optimal delay values used for He were 2–4 μs ; for air, they were 4–6 μs ; and for Ar, they were 12–14 μs . This result is in line with those known from the literature [16,17], according to which the electron temperature (T_e) and density (n_e) of the LIBS plasmas are inversely proportional to the atomic masses of the background gases [17], and a faster decay of both is observed for He.

It is also observed that the spectral emissions of atomic tungsten at 653.811 nm, 655.28 nm, 655.38 nm, 656.78 nm, 656.97 nm, and 657.18 nm are clearly detectable in the case of Ar and air plasma, while they are poorly detectable in the case of He plasma. This result agrees with previous results in the literature [18,19], the reason being that T_e and n_e decay more rapidly in He compared to Ar and air because He has a higher thermal conductivity than Ar and air [18,19].

To quantitatively estimate the resolution power, R_s , of the two D_{α} and H_{α} peaks in the case of the two gases He and Ar, the ratio between the width of the two peaks and their spectral separation was taken into consideration according to the following formula, which is currently applied in chromatography to evaluate the resolution of two nearby peaks:

$$R_s = \frac{2 \cdot \Delta\lambda}{(w_{D_{\alpha}} + w_{H_{\alpha}})} \quad (1)$$

where $\Delta\lambda$ is the spectral separation of the two peaks (0.18 nm), and $w_{D_{\alpha}}$ and $w_{H_{\alpha}}$ are the FWHMs of the two peaks, obtained by fitting them with two Gaussian functions.

The two peaks are resolved if $R_s \geq 1$. In Table 1 are the values obtained for the six spectra in Figure 1, acquired under He and Ar gas. Starting from these values, we estimated the resolution that would be obtained if tritium were present with the same concentration as deuterium, since the tritium T_α line (centered at 656.045 nm) is expected to have the same linewidth as the D_α line but with a spectral separation of $\Delta\lambda = 656.1 - 656.045 = 0.055$ nm. The theoretical resolution of the T_α and D_α peaks is reported in the last column of Table 1.

Table 1. Resolution of the two D_α and H_α peaks according to Formula (1) and theoretical resolution of the two T_α and D_α peaks with the same experimental conditions.

Background Gas	Delay (μ s)	R_s ($D_\alpha - H_\alpha$)	R_{s_theor} ($T_\alpha - D_\alpha$)
He	2	0.95	0.275
He	3	1.15	0.333
He	4	1.52	0.44
Ar	12	1.12	0.32
Ar	13	1.15	0.33
Ar	14	1.31	0.38

In any of the previous cases, the T_α and D_α peaks could not be resolved with the current system, and the discrimination between the two spectral signals should be performed through fitting procedures.

In this respect, a simulation of the LIBS signal in the presence of T was performed by adding to the experimental signal an additional, hypothetical T_α peak with the same FWHM and intensity as the experimental D_α peak, centered at 656.045 nm, which is the central wavelength of the T_α emission. In Figure 3, the results of the simulation (red curve) for the two most favorable cases are reported: He with a gate delay of 5 μ s and Ar with a gate delay of 14 μ s.

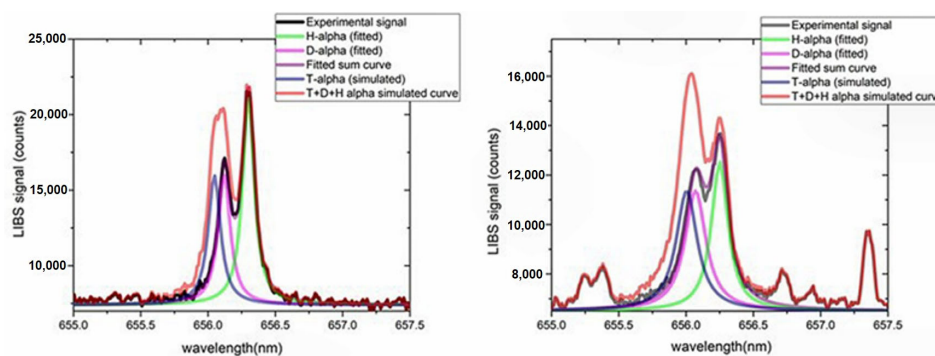


Figure 3. The simulated $T_\alpha + D_\alpha + H_\alpha$ spectra obtained by adding a T_α emission peak with the same width and height as the D_α peaks with He as a background gas (left) and Ar as a background gas (right). The experimental spectra were obtained with a gate delay of 5 μ s (He) and a gate delay of 14 μ s (Ar).

From the simulated spectra, the presence of an additional contribution at shorter wavelengths is evident in both cases in Figure 3 from the shape and width of the semi-theoretical signal, although the two peaks cannot be clearly resolved, as detailed in Table 1, and a post-acquisition fitting procedure would be necessary to separate the two components.

3.2. Depth Profiling Analysis

The LIBS depth profiling analysis [20] consists of consecutively hitting the sample at the same point with multiple laser pulses and recording the emitted spectrum of each pulse to gather information about the elemental composition with respect to depth. Each laser

pulse strips away a thin layer of sample, enabling the subsequent pulse to explore a slightly deeper portion.

In the case of the present measures, the depth profiling analysis was applied to the W-D sample to estimate the average ablation rate of the superficial W layer, co-deposited with D. For this type of analysis, the spectra acquired with the Echelle “full-range” spectrometer were used. The integral intensity, which was strong and relatively free of interference from the emission lines of W and Mo (W I at 400.87 nm and 407.44 nm and Mo I at 550.65 nm, 553.3 nm, and 557.04 nm, respectively), was calculated over a spectral range of 0.35 nm, centered around the peak wavelength and the (T)-D-H band (hereafter called T-D-H alpha), in the spectral window 656.0–656.3 nm. The integral intensities were monitored and reported as a function of the applied laser shots, as shown in Figure 4. A clear trend for the two W I lines and the T-D-H alpha band was observed, showing a higher integral intensity in the first 10 shots, then dropping sharply in the subsequent shots, demonstrating that the co-deposited W-D layer is present only on the surface of the sample.

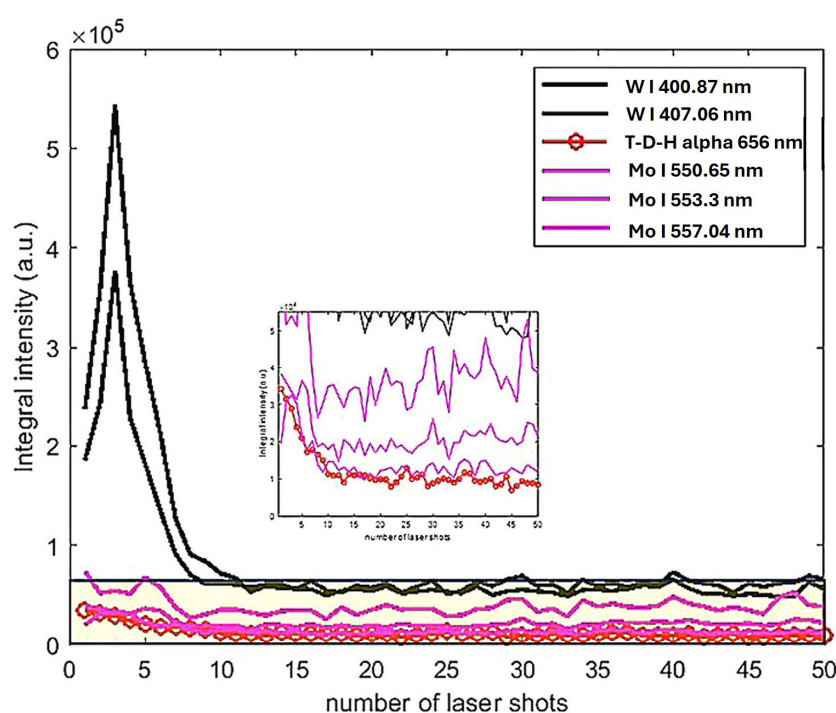


Figure 4. The integral intensities of the W I and Mo I lines, together with the band at 656.0–656.3 nm, as a function of the applied laser shots. The inset shows a magnification of the signal in the yellow rectangle, highlighting the signal related to the Mo I lines and the T-D-H band.

Assuming the nominal thickness of the superficial layer (5 μm), the average ablation rate of the W layer was estimated to be ≈ 500 nm/shot.

This result was also confirmed by the analysis through the non-contact optical profilometry of the laser crater obtained after applying 10 consecutive laser shots, as shown in Figure 5.

A z-scan, normal to the surface, was performed by the optical profilometer over an area of $550 \times 550 \mu\text{m}^2$, a vertical range of 30 μm , and a step of 10 nm to obtain the 3D reconstruction of the crater of Figure 5. The average depth of the crater was measured to be $\approx 4.73 \mu\text{m}$, in line with the result obtained by observing the trend of the emission intensity of the W and D + H lines.

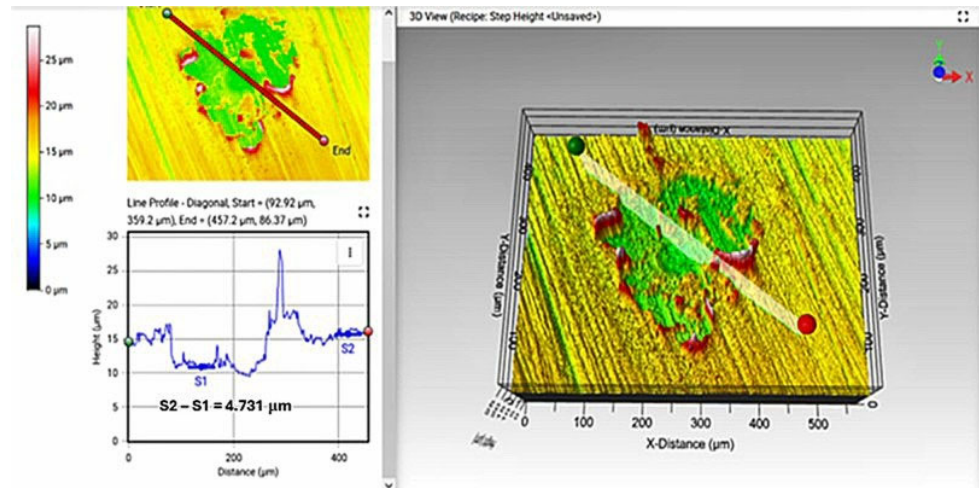


Figure 5. The optical profilometry of the LIBS crater measured after 10 laser shots. On the left are the color scale bar, the 2D image of the crater (up), and the depth profile along the line marked by the two red and green spheres (bottom). On the right is the 3D image with the step height between the two reference levels: S1 (average level of the bottom of the crater) and S2 (average level of the surface of the sample).

3.3. Calibration-Free Analysis

The CF technique is a methodology used for a quantitative estimation of the chemical species present in an LIBS spectrum by using the experimental line intensities. CF was applied to the LIBS spectrum obtained in Ar as a background gas with a gate width of 3 μs and a gate delay of 14 μs . The reason for applying CF to this spectrum is that well-resolved $D_{\alpha} - H_{\alpha}$ emission lines were detected, as well as W I emission lines at 653.811 nm and 657.393 nm; therefore, the quantitative estimation of D and H with respect to W is feasible.

In CF analysis [15,21,22], there is no need for referenced samples or calibration curves, but the laser-induced plasma is required to be optically thin [19,20] and in local thermodynamic equilibrium (LTE) [22]. In these conditions, the line integral intensities of the emission lines are related to the concentration of the species through the following expression:

$$I_{ki} = C_s A_{ki} g_k \frac{e^{-(E_k/k_B T)}}{U_s(T)} \quad (2)$$

where C_s is the species (atomic or ionic) concentration, A_{ki} is the transition probability for the specific emission line, g_k is the k-level degeneracy, k_B is the Boltzmann constant, T is the plasma temperature, and $U_s(T)$ the partition function of the species. To be applied, the CF analysis requires knowledge of T_e and n_e , which can be obtained from the experimental spectra through the Boltzmann plot [23–25] of the emission lines present in the spectra (T) and from the Stark broadening [26,27] of well resolved lines for which the Stark parameter is known [28] (n_e). The criterion for determining whether the LIBS plasma is optically thin is obtained from the intensity ratio of two lines of a particular element in the same charge state (ionic or atomic), which, according to Equation (2), is [29]

$$\frac{I_{ki1}}{I_{ki2}} = \frac{A_{ki1} g_{k1}}{A_{ki2} g_{k2}} e^{-\frac{(E_{k1}-E_{k2})}{k_B T}} \quad (3)$$

where the constants are the same as those described in Equation (2) for the two emission lines (labeled 1 and 2).

If the two emission lines either have the same or as close as possible upper levels, the contribution of the exponential factor in Equation (3) is minimized and can be neglected.

Therefore, the theoretical value of the intensity ratio of the two lines depends only on the spectroscopic parameters of the transitions and on the experimental intensity of the lines.

The optical thinness of the LIBS plasma was checked by applying Equation (3) to the intensity ratio of the two W I emission lines at 653.811 nm and 657.393 nm, whose spectroscopic parameters were retrieved from the NIST database [30]. These parameters, together with the theoretical and the experimental intensity ratios, are reported in Table 2:

Table 2. Spectroscopic parameters of the two W I emission lines at 653.811 nm and 657.393 nm and the relative theoretical and experimental ratios according to Equation (3).

λ (nm)	A_{ki} (10^{-8} s)	g_k	E_k (eV)	I_1/I_2 (Theo.)	I_1/I_2 (Exp.)
653.811	2.7×10^{-3}	9	4.354	0.975	1.09
657.393	9.9×10^{-3}	3	4.487		

From this table, it can be observed that the theoretical and experimental intensity ratios of the two lines are comparable, and the plasma can be considered optically thin.

To evaluate the LTE of the LIBS plasma, it is necessary to apply the so-called McWhirther criterion [31], which is a necessary but insufficient condition that can be expressed as follows:

$$n_e (\text{cm}^{-3}) > 1.6 \cdot 10^{12} T^{1/2} (\Delta E_{ki})^3 \quad (4)$$

where T (eV) is the plasma temperature and ΔE_{ki} (eV) is the maximum energy difference between the upper and lower energy levels. In the case of the sample under analysis made of W and H(D), we have ΔE_{ki} (eV) = 2.49 eV for W I, ΔE_{ki} (eV) = 4.71 eV for W II, and ΔE_{ki} (eV) = 10.2 eV for H(D). The plasma temperature T was retrieved from the extended Boltzmann plot [12,16,22] of the sequence of the 50 LIBS spectra applied to the same point and was found to be 8900 ± 400 K, whereas the electron density was evaluated through the evaluation of the Stark broadening of the three W I lines at 426.9 nm, 429.4 nm, and 430.2 nm [32] and was found to be $4.08 \pm 2 \times 10^{16}$ (cm^{-3}). Checking the McWhirther criterion for W and H, the condition is fulfilled.

In addition to the above criteria, the condition for the validity of LTE was also verified by calculating the diffusion length, δ (cm), through the following relation [22]:

$$\delta \approx 1.4 \times 10^{12} \frac{(kT)^{3/4}}{n_e} \left(\frac{\Delta E}{M_A \langle \bar{g} \rangle} \right)^{1/2} e^{\frac{\Delta E}{kT}} \quad (5)$$

and the atomic emission spectral line of W I at 251.814 nm was used to calculate δ , giving $\delta \approx 0.05$ mm. The characteristic variation length in the plasma, “ d ”, which can be taken as the plasma diameter (≈ 2 – 3 mm), is much larger than 10δ (i.e., $10\delta < d$), which ensures the second condition of LTE.

With both conditions satisfied (plasma optically thin and in LTE), the CF can be applied to estimate the relative concentration of H + D ([H + D]) with respect to W for all 50 spectra of the depth profiling analysis. The obtained values for [H + D] span from $0.5 \pm 0.5\%$ to $3.5 \pm 0.5\%$, due to uncertainties in the values of T and n_e , and are reported in Figure 6 as a function of the applied laser shots.

Looking at the trend of [H + D] values, it is observed that the surface layer appears enriched in hydrogen isotopes, decreasing until the 10th–15th shot, which is consistent with the qualitative results shown in Figure 4, where the H + D signal is above the level of noise within the first 15 shots. It is also observed that, after the 35th shot, the [H + D] values become noisier, with large variations. This is probably due to the complete ablation of the residual W layer at the edges of the crater (lateral ablation) and the consequent decrease in the W signal in the LIBS spectra, whereas the H + D signal remains constant in comparison

to a background level because of the constant fraction of environmental hydrogen ionized nearby the plasma plume.

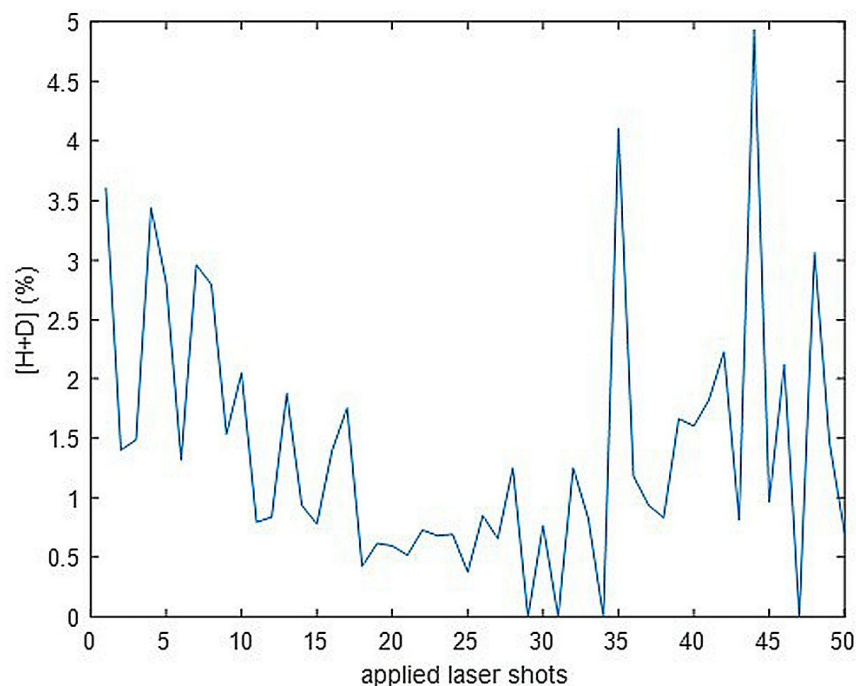


Figure 6. Relative (atomic) concentration of H + D hydrogen isotopes with respect to W as a function of the applied laser shots.

4. Conclusions

In this work, we presented the results of the application of the LIBS technique to a deuterated tungsten sample, simulating the surface of a PFC of a real tokamak after an experimental campaign with fusion fuel. The measurements were performed in He and Ar flows and in air to compare the possibility of resolving the spectral lines of hydrogen isotopes, which is of great interest for the application of the technique in the fusion field. The use of He and Ar as background gases allows for the separation of the H_{α} and D_{α} emission lines of hydrogen and deuterium with higher resolution than traditional LIBS in air, as soon as the acquisition parameters have been optimized. The simulation of the presence of tritium in the spectrum, with its T_{α} spectral emission, having an intensity comparable to the D_{α} emission of deuterium, allows us to conclude that the simultaneous presence of the two isotopes requires a system with a higher resolution for them to be clearly detected, and a post-measurement fitting procedure may be necessary to resolve the two lines. The depth profiling measurements highlighted an average ablation rate of about 500 nm per laser shot. Measurements in an Ar atmosphere also allowed us to observe the emission lines of W, which represents a material considered for PFCs in next-generation fusion devices, with a better signal-to-noise ratio compared to that in He atmosphere and comparable spectral resolution. Furthermore, measurements in Ar do not introduce an interfering signal if the analysis aims to reveal the He embedded into PFCs because of fusion reactions. Finally, the application of the CF technique allowed for a semi-quantitative estimation of the hydrogen isotopes co-deposited on the surface layer of the sample.

Author Contributions: Conceptualization, S.A.; methodology, S.A., M.P. and L.B.; validation, L.B.; investigation, S.A. and L.B.; data curation, S.A.; writing—original draft preparation, S.A.; writing—review and editing, S.A. and L.B. All authors have read and agreed to the published version of the manuscript.

Funding: This work was carried out within the framework of the EURO-fusion Consortium, funded by the European Union via the Euratom Research and Training Programme (Grant Agreement No 101052200—EUOfusion). The views and opinions expressed are, however, those of the author(s) only and do not necessarily reflect those of the European Union or the European Commission. Neither the European Union nor the European Commission can be held responsible for them.

Data Availability Statement: The data presented in this study are available on request from the corresponding author. The reason is due to the ongoing measurements and analyses for which the data are considered as reference.

Acknowledgments: The authors wish to acknowledge Eduard Grigore for the sample production Violeta Lazic and Francesco Colao for the design and settings of the LIBS system.

Conflicts of Interest: The authors declare no conflicts of interest.

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